Impact cratering, volcanism and tectonism on Solar system bodies

Impact cratering

Bolide: Photo of Hiroyuki-lida, Japan



Sterlitamak impact crater, Russia



Meteor crater, panorama of crater interior



Meteor crater rim seen from the road



Roter Kamm crater, Namibia, D = 2.5 km



New Quebec crater, Canada, D = 3.4 km



Elgygytgyn crater, Chukotka, Russia, D = 18 km



Elgygytgyn crater, Dr. Feldman is exploring it



Impact crater Nordlingen Ries, Germany



D = 23 km

Bunte (pied) breccia, crater Ries, Germany



Suevute, crater Ries, Germany



Moldavite: splash of impact melt from crater Ries



Craters Clear Water West (36 km) & East (26 km) Quebec, Canada



Popigai crater, Siberia, Russia, D = 100 km



Popigai crater, geologic map of the region



ГЕОЛОГИЧЕСКАЯ КАРТА ПОПИГАЙСКОЙ АСТРОБЛЕМЫ

Масштаб 1:1 000 000

Составили: В.Л.Масайтис, М.С.Мащак, М.В.Михайлов, А.Н.Данилин, В.Т.Кириченко

1978



A

M 10007

-1000-

-2000-

-3000-

-4000-

Popigai geologic map by Masaitis et al., 1978

дополнительные условные обозначения





-2000

-3000

-4000

Апогнейсовая аутигенная брекчия

Министерство Всесоюзный ордена Ленин геологический и Лени

paspesa

ялд

Только



Popigai crater (astrobleme) Megabreccia and tagamites at Rassokha river



Tagamite = solidified impact melt, crater Janisjarvi, Russia



Sudbury crater (astrobleme) Has very rich nickel deposit inside



Sudbury crater geologic map



Sudbury shutter cones



Vredefort crater, South Africa, initially D ~ 300 km



Formed

2 b.y.

ago

Has gold deposits within

Tunguska event



Tunguska event: radially fallen forest



World map of impact structures



Koeberl: http://www.lpi.usra.edu/publications/slidesets/craters/

Lunar simple crater



Lunar complex crater: central-peaked



Lunar complex crater: double-ringed



Lunar Orientale basin: multi-ringed



Morphology of impact craters depends on their size от их диаметра

On the Moon <10 km – bowl-shaped 15-20 km – transition to craters with central peak 20-90 km – craters with central peak > 90 km – ringed basins





Lunar crater Copernicus with secondary crater around

Mercury

Small craters – bowl-shaped Large – with central peak Very large – ringed basins



Messenger, NASA



Mariner 10 images

Impact craters of Venus



Impact craters of Mars: Small crater bowl-shaped, large - with central peak and excavate material, containing water ice



with «normal» ejecta



Crater Yuti with central peak, ejecta are fluidized (mud flows)

100-km crater with central peak on Mimas (D = 400 km)

Impact basin on Callisto (D = 4680 km)

300 км

Voyager 2 image

Galileo image
Highland breccia Apollo 16, Fragments are well seen



Grains of quartz affected and not affected by impact

20 micron



No impact effects

Planar structures in impacted quartz

http://www.lpi.usra.edu/publications/slidesets/craters/



Planar structures in impacted quartz Crater Janisjarvi, Karelia, photo of L.V. Sazonova

Impact diamonds from Australia and Brazil



Impact diamond. Quinsland, Australia

Impact diamond (carbonado) Brazil

Stages



of impact

http://www.psi.edu/explorecraters/ intro_pics/impact_stages_16.jpg

Structure of impact craters: Synthesis of terrestrial and lunar data



Mean speed of collision (km/s) of meteoroids with planetary bodies

Planetary body	Escape velocity	Early planatesimals	Asteroids	Comets
Mercury	4.3	4.7	20	62
Venus	10.3	11.5	18	47
Earth	11.2	12.5	18	40
The Moon	2.37	6.1	14	38
Mars	5.03	5.6	10	31

Specific kinetic energy of projectile (erg/g) in comparioson with specific energy of explosion

TNTNuclearV = 5 km/sV = 15 km/sV = 30 km/sV = 45 km/s 4×10^{10} 1×10^{16} 1.2×10^{11} 1.1×10^{12} 4.5×10^{12} 1×10^{13}

Impact induced transformations of materials 1 bar = 1 km/cm2, 1 kbar = 103 bar, 1 Mbar = 106 bar

Mechanical crushingSeveral kbarModification of crystal structure100-200 kbarPlanar elements100-200 kbarIsotropization of minerals SiO2260-300 kbarTransformation into glass with preservation of rock structure

M impact vapor \approx M projectile M impact melt \approx 10 M projectile

Experimental data on energy distribution (%) in impacts

Target – sand Target - basalt Heating Projectile ~6 4-12 19-23 ~25 Target Compaction ~20 1 Crushing ~8 10-24 Ejection ~53 42-53

For Earth D projectile $\approx 1/10 - 1/20$ D crater Crater with D = 10 km was formed by ~1 km projectile

Chicxulub crater (astrobleme) location

© 2000 by Jake Bailey Adapted from "Atlas of Mesozoic and Cenozoic Coastlines" (Smith et al. 1994)

Chicxulub gravity map



Structure of crater Chicxulub



K-T boundary, Raton basin, Colorado, USA



Impact which killed dinosaurs, art by Don Davis





Origin of the Moon: Giant impact hypothesis

Mars-size body



The Moon is formed due to accretion of the impact ejecta on the around-Earth orbit



Magmatism (volcanism) on the Solar system bodies



Plate tectonics -- working heat engine of Earth

How plate tectonic works

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Sea-floor spreading lava eruption: Pillow lavas



Several meters

Sea-floor spreading: small volcanic ridge at Mid Atlantic Ridge / _{Sonar image}



East African rifts





Nyiragongo, East Africa



Hawaii big island: hot spot volcanism



Landsat image

Map

Hawaii, Mauna Loa summit



Caldera Kilauea eruption

Mauna Loa eruption



Hawaii, aa and pahoehoe lavas



Lava fountaining forms cinder cones, Hawaii



Cinder cone and Mauna Kea volcano behind



St. Helens volcano: pre eruption



St. Helens blast and nuee ardente May 18, 1980



Blast

Nuee ardente

USGS photo

St. Helens pyroclastic tongue



St. Helens pyroclastic deposit



St. Helens dome, August 1981



Lunar maria

Basaltic lavas



Lunar volcanism



Plains-forming basaltic lavas make maria.

Sinuous rilles – erosion by hot lava flow


108 km crater Alphonsus has on its floor fractures and small craters with dark halo – pyroclastic deposit?



In 1958 Soviet astronomer N. Kozyrev observed here a gas release

Gruithuisen domes -- non-basaltic volcanism?



Mercury: Smooth plains resembling lunar maria



Volcano in Caloris basin, Mercury

- On volcano are seen irregular rimless depressions
- There are rather large impact craters aroun the volcano
- On the volcano surface are seen onle small craters => large difference in time
- Source of the lava is not the impact melt





Venus: Plains made of basaltic lavas. Radar bright lavas are probably close to the aa type.

Venus: Shield volcanic plains



Volcanic shields with gentle (3-5 deg) slopes Probably formed by basaltic lavas

Lobate lava plains



Volcanic flows with rather rough (aa) surface, Superposed on plains with wrinkle ridges and shield plains

Maat Mons volcano – highest on Venus (+9 km)



Maat slopes are covered by radar bright lobate flows superposed on surrounding regional volcanic plains

Maat Mons volcano – perspective view



Vertical exaggeration ~ 20:1

Rare type of volcanic constructs on Venus: steep-sided domes



Tens km in diameter, Hundreds meters high, Steep-sided: => Viscous lavas => Evolved composition e.g. dacites? or => Basalt with gas bubbles?



Mars: Olympus Mons, 600 km across, 21 km high





Intermediate and small volcanoes of Mars



Ceraunius Tholus

50 km

Courtesy of Calvin J. Hamilto

Cinder cones on the surface of Mars





Volcanism on Io: closest to Jupiter Galilean satellite of the planet

Distance to center of Jupiter ~6 R Jup.

Image taken by Galileo

Tidal heating Plumes of volcanic eruptions Orange color – sulfur But lavas are basaltic No impact craters => very young surface

The eruption plume height is 160 km

Voyager 1 image

Io: Chain of volcanic calderas Twashtar



Europa – second Galilean satellite of Jupiter – faults, domes and reddish spots and bands



Multiphase tectonics / water-ice (cryo) volcanism



Enceladus: Soth pole Geisers of H₂O vapor => Result of tidal heating

Area of source of geiser #6

BAGHDA SULCUS 5 km Cassini 31.10.2008

Triton – satellite of Neptune Plains – products of water-ice (cryo) volcanism



200 km

https://www.nasa.gov/image-feature/mosaic-of-high-resolution-images-of-plute

Nitrogen cryovolcanism on Pluto

Sputnik Planum

Tectonism

Spreading: The Read Sea rifts and Aqaba Bay



Extension structures: faults and graben





Faults



Graben

Subduction zone: Compresion: Cascades



Compression structures: folds









Transform faults St. Andres, California



Lunar graben – structures of extension



Lunar sinuous rilles: structures of compression



Tectonism of Mercury:

- Ridges and graben
- Long scarps (upthrusts) up to 1-3 km high

=> Compression of planet with decrease of radius by 2 km



Scarps

Graben

Rift zones of Venus



Beta Regio – tectonic uplift, cut by Devan rift



Devana rift and impact crater Balch



Lakshmi plateau and mountain belts around



Venus: Tesserae: multiphase tectonism Compression and extension



Venus: Ridge belts: Compression



Tectonism of Mars: Mostly extension, partly compression



Valles Marinares – Rift zones
Mars: Graben parallel to Valles Marinares



Mars: Sinuous rilles indicate on compression



In morphology and size are close to those of the Moon and Venus

Europa: Faults and folds



Faults on Ganimede



Titania (D = 1580 km) – satellite of Uranus



Triton – satellite of Neptune Tectonic deformations in the cantalupa terrain



Presence / intensity of magmatic and tectonic activity on planets / satellite is a function of their sizes and tidal disturbances by the central body



Impact processes were and are in work, but their traces are well seen only where other geologic processes are weak

Sputnik 1 launch, October 4, 1957

This image was composed by Don Mitchell from the nresreel, Don Mitchell, www.mentallandscape.com